ZARM, University of Bremen

July 28, 2010

Towards a derivation of G

Frans R. Klinkhamer

Institute for Theoretical Physics, University of Karlsruhe, Karlsruhe Institute of Technology Email: frans.klinkhamer@kit.edu

a somewhat unusual talk ...

ZARM, Bremen, July 28, 2010 (v1) - p. 1

Physical constants

Fundamental physical constants (Roemer, Cavendish, Planck):

$$c, \quad G, \quad \hbar$$

Theorists often set c = 1, G = 1, and $\hbar = 1$, by using appropriate units for length, time, and energy.

This practice considers SR (= special relativity), GR (= general relativity), and QM (= quantum mechanics), to be closed chapters.

But what does Nature say?

Physical constants

Table 1: Known constants of nature [1].

quantum matter	classical relativity	quantum spacetime
(Planck & Bohr)	(Einstein)	(Wheeler)
ħ	c,G	$l_P \equiv \sqrt{\hbar G/c^3}$

Possible argument for a single constant \hbar controlling the quantum nature of <u>both</u> matter (e.g., photons & electrons) and spacetime:

- quantized electrons ↔ quantized electromagnetic field ⇒ QED [exps: Geiger and Bothe, 1925; Compton and Simon, 1925];
- similarly, quantized electrons \leftrightarrow quantized metric field? \Rightarrow ???

^[1] P.J. Mohr, B.N. Taylor, and D.B. Newell, RMP 80, 633 (2008), arXiv:0801.0028.

Physical constants

Table 2:Alternative constants of nature [2].

quantum matter	classical relativity	quantum space
ħ	$c,G\equiv fc^3l^2/\hbar$	l^2

Possible arguments for a <u>new</u> constant l^2 of quantized space:

- space (and gravity) may be emergent phenomena;
- <u>natural</u> to have a constant with dimension of length/area/volume.

Conceptual remark:

If Table 2 holds true, there may be physical situations where matter quantum effects are negligible (" $\hbar = 0$ ") but not spacetime quantum effects (" $l^2 \neq 0$ "), which is impossible if Table 1 holds ($l_P = 0$ for $\hbar = 0$).

[2] F.R. Klinkhamer, JETPL 86, 73 (2007), arXiv:gr-qc/0703009.

Now, consider **Newtonian gravity**, specifically, the inward acceleration \vec{A}_{grav} on a test mass *m* produced by a point mass *M* at a distance *R*:

$$\vec{A}_{\rm grav} = -(GM/R^2)\,\hat{e}_R\,. \tag{1}$$



Newton (1713): "hypotheses non fingo"

Using the G formula from Table 2,

$$G = f c^3 l^2 / \hbar , \qquad (2)$$

with a factor f > 0, the magnitude of this acceleration reads [2]

$$A_{\rm grav} = GM/R^2 = f c \left(M c^2 / \hbar \right) \left(l^2 / R^2 \right), \tag{3}$$

where all microscopic quantities are indicated by lower-case symbols.

Possible interpretation of the two factors in brackets on the RHS of (3):

- first factor is a decay rate of space triggered by external mass M;
- second factor is a *geometric dilution factor*.

Interpretation perhaps suggestive but definitely vague.

Progress from a recent idea of Verlinde ...

•

Entropic gravity

Verlinde's proposal [3] is that Newtonian gravity arises as an **entropic force** from a **holographic** [4] microscopic theory.

Main steps [3]:

- holographic screen $\Sigma(x_1, x_2)$ with orthogonal dimension x_3 emerging from coarse-graining degrees of freedom (d.o.f.) on Σ ;
- entropy change from nearby mass m at distance Δx_3 is given by $\Delta S_{\Sigma} \propto (m c/\hbar) \Delta x_3$;
- first law of thermodynamics: $T\Delta S_{\Sigma} = F_{\text{grav}}\Delta x_3 \Rightarrow F_{\text{grav}} \propto mM/R^2$, with mass equivalent M of spherical screen with area $4\pi R^2$.
- [3] E. Verlinde, arXiv:1001.0785v1.
- [4] G. 't Hooft, arXiv:gr-qc/9310026; L. Susskind, JMP 36, 6377 (1995), arXiv:hep-th/9409089.

m



Left panel: Spherical holographic screen Σ_{sph} with area $A = 4\pi R^2$ and test mass m. Space has emerged outside the screen Σ_{sph} , which has N microscopic degrees of freedom at an equilibrium temperature T with total equipartition energy $E = \frac{1}{2} N k_B T$. <u>Right panel</u>: The gravitational effects of Σ_{sph} for the emergent space correspond, in leading order, to those of a point mass $M = E/c^2$ located at the center of a sphere with radius R (the Schwarzschild radius $R_{Schw} \equiv 2GM/c^2$ taken negligible compared to R).

With this spherical holographic screen Σ_{sph} [3], a different 'derivation' [5] may give a clue to the origin of the previous 'suggestive' formula (3):

$$A_{\text{grav}} \stackrel{\textcircled{1}}{=} 2\pi c \left(k_B T / \hbar \right)$$

$$\stackrel{\textcircled{2}}{=} 4\pi f c \left(\frac{1}{2} N k_B T / \hbar \right) \left(f^{-1} / N \right)$$

$$\stackrel{\textcircled{3}}{=} 4\pi f c \left(E / \hbar \right) \left(l^2 / A \right)$$

$$\stackrel{\textcircled{4}}{=} f c \left(M c^2 / \hbar \right) \left(l^2 / R^2 \right), \qquad (4)$$

where step 1 relies on the <u>Unruh effect</u> [6] and step 3 on the relation between the number N of d.o.f. and the area A of the holographic screen:

$$N = f^{-1} A/l^2 \,. \tag{5}$$

[5] F.R. Klinkhamer, arXiv:1006.2094v3.[6] W.G. Unruh, PRD 14, 870 (1976).

The several steps of (4) constitute, if confirmed, a **derivation** of Newton's gravitational coupling constant G in the form (2).

New insight from (5): given the "effective quantum of area" l^2 , the <u>inverse</u> of the constant f entering Newton's constant (2) may be related to the <u>nature</u> of the microscopic d.o.f. on the holographic screen.

For example, an "atom of space" with "spin" s_{atom} may give $f^{-1} = 2 s_{\text{atom}} + 1 \equiv d_{\text{atom}}$, but s_{atom} need not be half-integer.

Therefore, rewrite (5) as

$$N = d_{\text{atom}} N_{\text{atom}}, \quad d_{\text{atom}} \equiv f^{-1} \in \mathbb{R}^+ \quad N_{\text{atom}} \equiv A/l^2 \in \mathbb{N}_1, \quad (6)$$

where the "atoms of space" (total number N_{atom}) have no translational degrees of freedom but only internal degrees of freedom (d_{atom}).

Next, calculate the factor $f \equiv (d_{atom})^{-1}$ entering formula (2) for G.

Consider a maximally-coarse-grained spherical surface (horizon) with area A. Entropy given by the Bekenstein–Hawking black-hole result [7]:

$$S_{\text{BH}}/k_B = (1/4) A/(f l^2) = (1/4) N.$$
 (7)

Equating the number of configurations of the "atoms of space" from (6) with the exponential of the BH entropy (7) gives a **set of conditions** [5]:

$$(d_{\text{atom}})^{N_{\text{atom}}} = e^{(1/4) \, d_{\text{atom}} \, N_{\text{atom}}} \,, \tag{8}$$

which reduces to a single transcendental equation for d_{atom} :

$$4 \ln d_{\text{atom}} = d_{\text{atom}} \,. \tag{9}$$

This equation has two solutions:

$$d_{\rm atom}^{(+)} \approx 8.613\,169\,456\,, \quad d_{\rm atom}^{(-)} \approx 1.429\,611\,825\,.$$
 (10)

[7] J.D. Bekenstein, PRD 7, 2333 (1973); S.W. Hawking, CMP 43, 199 (1975).

Given l^2 , there are then two possible values for the gravitational coupling constant (2):

$$G_{\pm} = \left(d_{\text{atom}}^{(\pm)} \right)^{-1} c^3 l^2 / \hbar \,. \tag{11}$$

The detailed microscopic theory must tell which of the two d_{atom} values from (10) enters (11).

It could, for example, be that the microscopic theory demands $d_{\text{atom}} \ge 2$, selecting the value $d_{\text{atom}}^{(+)} \approx 8.6$ and giving

$$G_{+} \approx \left(8.613\,169\,456\right)^{-1} c^{3}\,l^{2}/\hbar \approx \left(0.116\,101\,280\,1\right) \,c^{3}\,l^{2}/\hbar \,.$$
 (12)

But the experimental value of Newton's gravitational coupling constant is already known (to 100 ppm [1]): $G_N = 6.6743(7) \ 10^{-11} \ \text{m}^3 \ \text{kg}^{-1} \ \text{s}^{-2}$.

A more practical interpretation of result (10) for d_{atom} is, therefore, to calculate two possible values for the "effective quantum of area":

$$(l_{\pm})^2 = d_{\text{atom}}^{(\pm)} (l_P)^2 \approx \begin{cases} 2.2498 \times 10^{-69} \text{ m}^2, \\ 3.7343 \times 10^{-70} \text{ m}^2, \end{cases}$$
(13)

with $l_P \equiv (\hbar G_N)^{1/2} / c^{3/2} \approx 1.6162 \times 10^{-35}$ m.

The microscopic theory would, again, have to choose between these alternative values.

For either choice, the implication would be that l and l_P are of the same order of magnitude.

The crucial question, now, is if l^2 can be measured directly.

Possible experiments:

- Cosmic-ray particle-propagation experiments (e.g., Auger) can search for Lorentz-violating effects from a nontrivial small-scale structure of spacetime [2] and may determine the ratio $f = (l_P/l)^2$ if the size of spacetime defects is set by l_P and their separation by l.
- A Gedankenexperiment can measure quantum modifications [8] of Newton's gravitational acceleration (3) by a multiplicative factor $\left[1 \tilde{a} l^2/R^2\right]$ and determine l^2 if $\tilde{a} > 0$ is known from theory.

But the if's make these experiments inconclusive, for the moment.

Perhaps further examples from the conceptual remark below Table 2.

^[8] L. Modesto and A. Randono, arXiv:1003.1998v1.

Finally, two remarks on the numerical value of G_N .

First, the order of magnitude is given by (using mks units):

$$G_N \sim c^3 \, \frac{l^2}{\hbar} \sim \left(3 \times 10^8\right)^3 \, \frac{3 \times 10^{-70}}{1 \times 10^{-34}} \sim 10^{-10} \, \mathrm{m}^3 \, \mathrm{kg}^{-1} \, \mathrm{s}^{-2} \,, \quad \text{(14)}$$

Second, note that the accurate measurement of one of the values of l^2 in (13) allows for an equally accurate calculation of G from (11).

For example, measuring for l^2 the larger value in (13) with a relative uncertainty of 100 ppb would give *G* also with an uncertainty of approximately 100 ppb from (12):

$$G_N \stackrel{?}{=} G_+ \approx \left(0.116\,101\,280\,1\right) \,c^3 \,l^2/\hbar\,.$$
 (15a)

If, instead, the smaller value for l^2 would be measured to 100 ppb, then

$$G_N \stackrel{?}{=} G_- \approx \left(0.699\,490\,576\,9\right) \,c^3 \,l^2/\hbar\,.$$
 (15b)

Conclusion

Two interesting results:

- derivation' of $G = f c^3 l^2 / \hbar$ via Unruh temperature & holography;
- calculation of $f \equiv (d_{atom})^{-1} = (l_P/l)^2$ from BH black-hole entropy.

Many outstanding questions:

- Are space and gravity really emergent phenomena?
- If so, really from a holographic theory?
- Also, is Newton's gravitational force really an entropic force ?
- Independently, is there a <u>new fundamental constant l^2 ?</u>
- If so, what is the value of f in the relation $G = f c^3 l^2 / \hbar$?
- Also, how can l^2 be <u>measured</u>, in principle and in practice?